

Constraining cosmological parameters in FLRW metric with lensed GW+EM signals

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GW150914 B. P. Abbott et al. 2016



EM



EM+GW











GW

- Ding et al.(2015): ET will detect 50 100 lensed GWs per year
- 2. Liao et al.(2017): just 10 such systems can provide a Hubble constant uncertainty of 0.68%
- 3. Fan et al.(2017): lensed GW+EM systems can be used to measure the speed of gravitational waves
- 4. Wei et al.(2017): strongly lensed GW-EM systems can provide stringent constraints on parameters of Dark energy equation of state



Lens Source redshift redshift $D_{\Delta t} = \frac{D_A(z_l)D_A(z_s)}{D_A(z_l, z_s)}$ $D_{\Delta t}(theory) = \frac{1}{(1+z_i)(T(z_i) - T(z_c))}$ $D_{\Delta t}(observation) = \frac{\Delta t}{(1+z_l)\Delta \phi}$ $\Delta \phi_{i,j} = \left[\frac{(\theta_i - \beta)^2}{2} - \psi(\theta_i) - \frac{(\theta_j - \beta)^2}{2} + \psi(\theta_j)\right]$ $T(z) = \frac{1}{d(z)} \sqrt{1 - kd(z)^2}$ $d(z) = z + a1z^2 + a2z^3 + a3z^4$ Fermat potential difference $D_{\Delta t}$ (observation) = $D_{\Delta t}$ (theory) observed quantities (z_l , z_s , Δt , $\Delta \varphi$, D_L) theoretical values (k, d(z), H0)





https://arxiv.org/abs/1901.10638

	No.	<i>a</i> 1	a2	<i>a</i> 3	k	H_0
lensed quasar	300	$-0.24^{+0.04}_{-0.04}$	$0^{+0.06}_{-0.06}$	$0.01\substack{+0.03 \\ -0.02}$	$0^{+0.05}_{-0.06}$	$69.86_{-0.53}^{+0.53}$
lensed GW+EM	10	$-0.25^{+0.02}_{-0.02}$	$0.02^{+0.02}_{-0.02}$	$0.01\substack{+0.01 \\ -0.01}$	$0.01\substack{+0.05 \\ -0.05}$	$69.7_{-0.35}^{+0.35}$

Summary

• Strongly lensed GW-EM systems combine highly accurate time delay from GW signals with redshift and image from EM counterparts.

- In the future, strongly lensed GW-EM systems can be applied to do precision cosmology.
- Establish a complete set of lensed GW-EM numerical simulation is necessary.
 Thank